Time Irreversible Billiards with Piecewise-Straight Trajectories

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(Received 27 June 2012; published 23 October 2012)

We introduce a new family of billiards which break time reversal symmetry in spite of having piecewise straight trajectories. We show that our billiards preserve the ergodic and mixing properties of conventional billiards while they may turn into exponential the power law decay of correlations characteristic of Sinai type billiards. Such billiards can be implemented by squeezing the transverse magnetic field along lines or along one-dimensional manifolds.

Introduction.—Billiards are mathematical models where one or more particles move freely in a container and collide with its walls and/or with each other [1–3]. The dynamical properties of such models, which play a basic role in the understanding of a large variety of physical phenomena, are determined by the shape of the walls of the container. Their behavior is very rich and can range from completely integrable, regular motion, via generic mixed phase space dynamics, to fully chaotic motion. Classical and quantum billiards play a basic role in nonlinear physics since in several cases they allow a rigorous theoretical analysis and also because they are very convenient for numerical simulation. They have been used in relation to the celebrated Boltzmann hypothesis in statistical mechanics, the Riemann hypothesis, number theory, the Lorentz gas, introduced in 1905 to describe electricity, and finally in applications related to low-dimensional nanostructures such as quantum dots [6], which are studied in view of emerging applications, including quantum computation. They exhibit deterministic ballistic motion of the electrons and provide the possibility to tune their shape, size, and electron number. Billiards have furthermore been experimentally realized for investigations in a variety of fields: in room acoustics and elastomechanics of solid plates, in atom optics, where ultracold atoms reflect from laser beams, in optics of dielectric microresonators, etc. Understanding the details of classical billiard dynamics even had a very practical application in controlling directed emission in solid-state microlasers [7].

From the theoretical point of view and for physical applications it is important to have billiards with broken time reversal symmetry. There are no examples of such models so far. The closest models are billiards in perpendicular magnetic field which have been shown to exhibit quantum interference effects like weak localization, Altshuler-Aronov-Spivak oscillations, and conductance fluctuations [8]. However, the dynamical properties of such models are much more complicated. In particular, with the exception of billiards with strictly convex boundaries [9], the application of a uniform magnetic field to an ergodic billiard leads to the appearance of islands of stability in phase space and to a very complicated dynamics. Indeed, the presence of the magnetic field changes the stability properties of the billiard motion. The Jacobian matrix of the motion between collisions, which is parabolic in the case without magnetic field, becomes elliptic. Moreover, the magnetic field changes also the parabolic matrix corresponding to boundary collisions. As a result, the study of the stability properties of the orbits becomes very difficult [10].

In this Letter we introduce a new class of billiards—i.e., conservative (Hamiltonian) systems in which motion takes place on straight lines apart from collisions with the boundary—which are time reversible symmetry breaking and do not alter the ergodic and mixing property of conventional billiards. It is remarkable that the decay of correlations can be turned from power law to exponential by just tuning the intensity of the magnetic field. These billiards may become important tools in investigation in several fields of nanoscience.

Billiards with magnetic boundaries.—The key ingredient in our idea is a one dimensional magnetic manifold, a strip of a transverse magnetic field which we shall
eventually think of squeezing into a line. Imagine such a locally flat strip of width \( d \), onto which a charged particle impacts under angle \( \theta \) (see Fig. 1). Then, depending on the strength of the magnetic field \( B \) in the strip, yielding a Larmour radius \( r = B^{-1} \) (we shall use units where the particle mass, charge, and velocity \( m = e = v = 1 \)), the particle either penetrates the magnetic strip, if \( |\sin \theta - d/r| < 1 \), or, oppositely, is reflected. Without losing generality we may assume that the magnetic field is pointing upward (from the plane of the figure), i.e., \( B > 0 \).

Introducing the nondimensional magnetic parameter

\[
\eta := \frac{d}{r} = Bd,
\]

one sees that the angle of incidence \( \theta \) and the angle of emittance \( \theta' \) are related by a trivial trigonometric identity (see Fig. 1):

\[
\sin \theta + \sin \theta' = \eta.
\]

This geometric relation remains valid in the limit where the magnetic strip becomes infinitely narrow \( d \to 0 \) if at the same time the strength of the magnetic field is increased \( B \to \infty \), such that \( \eta \) remains finite. The penetration or collision rule for such magnetic boundary then becomes

\[
\theta' = \begin{cases} 
\arcsin(\eta - \sin \theta), & \text{if } |\eta - \sin \theta| \leq 1; \\
-\theta, & \text{if } |\eta - \sin \theta| > 1.
\end{cases}
\]

Notice that, in spite of the fact that motion always occurs along straight lines, the magnetic boundary manifestly breaks time reversal symmetry: the time-reversed trajectory is in general not a valid trajectory. The nontrivial range of the magnetic parameter is \( 0 < \eta < 2 \). For \( \eta \geq 2 \), all trajectories get specularly reflected as in conventional billiards, while for \( \eta = 0 \) the walls are simply transparent.

Considering orbit stability, the magnetic boundary behaves in exactly the same way as a hard-wall boundary. In particular, if the magnetic boundary is locally flat, it maps parallel trajectories to parallel trajectories. More generally, it preserves the spread of the sines of the angle of a beam of trajectories, since \( d \sin \theta = -d \sin \theta' \), where \( d \) denotes the differential. If one uses appropriate canonical Birkhoff coordinates for the bounce map one can thus show that flat magnetic boundaries cannot change the stability focusing or defocusing property of the billiard. Since the corresponding Poincaré map is symplectic the uniform measure is an invariant measure of the dynamics. As a generic example, we study in the following the Sinai billiard with a magnetic outer boundary as indicated in Fig. 2. As the particles will in general be able to penetrate through the outer boundary (a unit square), we have to impose periodic boundary conditions in order to keep the phase space finite. However, the dynamics of such a billiard on the unfolded configuration space may be interpreted as the dynamics of charged particles in a magnetic heterostructure where the magnetic field is squeezed onto a Cartesian net with circular hard-wall obstacles of radius \( \rho \) placed in the center of each cell defined by the net.

As mentioned above, the standard proof of ergodicity and mixing for dispersing (Sinai-type) billiards [1], goes through along the same lines if pieces of flat boundary are replaced by our magnetic boundary. Thus our billiard should be ergodic and mixing with the positive Lyapounov exponent independently on the value of \( \eta \). In Fig. 3 we plot the Poincaré surface of a section, with respect to one side of the square (which is arbitrary due to the \( C_4 \) symmetry) of a Sinai billiard with magnetic boundary, and compare it to a surface of a section of the same Sinai billiard with a uniform magnetic field of comparable average strength. We note that, as expected, in the second case there are

**FIG. 1.** Schematic illustration of time-irreversible collision rules. The light-grey region of thickness \( d \) is considered to be in a transverse magnetic field where the Larmour radius is \( r \). At the end we consider the limits \( d \to 0 \) and \( r \to \infty \), such that the ratio \( \eta = d/r \) is finite. On the left we show penetration, which is possible when \( |d/r - \sin \theta| < 1 \), while in the opposite case we have a specular reflection (shown on the right).

**FIG. 2 (color online).** Schematic picture of the time-irreversible Sinai billiard. The boundary of a unit square (in light gray) is considered to be magnetic, while the circular obstacle of radius \( \rho \) is considered to be a hard-wall boundary. When the particle penetrates through the magnetic boundary we consider the periodic boundary condition and inject it on the opposite side of the square. Alternatively, we may consider a magnetic heterostructure (shown in the right) where the magnetic field is squeezed onto a two-dimensional Cartesian net.
islands of regular motion in sharp contrast to the ergodic behavior of our magnetic billiard. For the latter we have made careful numerical investigations and all evidence show that there are no islands. In Fig. 3 we just show one example to compare with the case of a uniform magnetic field. The time-irreversible nature of our billiard leads to the appearance of interesting dynamical properties. For example, a characteristic feature of Sinai-like billiards is the presence of the so-called long time tails. These tails are caused by the existence of a continuous family—of zero measure—of marginally stable periodic orbits which lead to power-law decay of correlations in spite of the mixing and exponentially unstable motion. Our magnetic billiard exhibits a richer behavior. In particular, by changing the intensity of the applied field one can control the decay of correlations without altering the general ergodic properties of the system. Indeed, upon changing \( \eta \), marginally stable periodic orbits may disappear, thus restoring the exponential decay of correlations. However, for \( \eta > 1 \), marginally stable periodic orbits do exist and are the same bouncing balls orbits (\( \theta = 0 \)) of the conventional Sinai billiard. Such orbits are also present for \( \eta < 1 \) provided the disc radius \( \rho < \rho_{cr} = \sqrt{2}/4 \). This restriction in the disc radius is necessary to allow the presence of diagonal (diamond shaped) marginally stable periodic orbits (\( \theta = \pi/4 \)) connecting the mid points of the squares sides. However, for \( 0 < \eta < 1 \) and for \( \rho > \rho_{cr} \), no marginally stable periodic orbits exists and, therefore, one expects exponential decay of correlations. These expectations are corroborated by numerical results shown in Fig. 4 in semi-log and log-log scale. Here we consider the Poincaré map (with respect to all four sides of the square, or with respect to the magnetic net of the 2D periodic heterostructure), and compute the velocity-velocity autocorrelation function,

\[
C(n) = \langle v_x(0)v_x(n) \rangle = \lim_{M \to \infty} \frac{1}{M} \sum_{m=1}^{M} v_x(m)v_x(m + n). \quad (4)
\]

FIG. 3 (color online). Poincaré surface of a section for the Sinai billiard with magnetic boundary (\( \eta = 0.5 \)) [top panel] and for the Sinai billiard with an all-hard-wall boundary and a uniform magnetic field with Larmour radius \( r = 0.5 \) [bottom panel]. In both cases the discs’ radius is \( \rho = 0.3 \). We note the appearance of islands of regular motion in the second case, while in the first case the motion is ergodic. Here we use only one side of the unit square as a surface of the section (0 \( \leq x \leq 1 \)) with the canonically conjugate momentum coordinate \( \sin \theta \).

FIG. 4 (color online). Velocity-velocity correlation function \( C(n) = \langle v_x(0)v_x(n) \rangle \), where \( n \) counts the number of collisions (penetrations) with (through) the magnetic boundary. We take \( \rho = 0.4 \), and consider different values of the magnetic parameter \( \eta = (\sqrt{5} - 1)/2 \) (black circles), \( \eta = 0.1 \) (blue squares), for which we find exponential decay \( \propto \exp(-0.45n) \) (indicated by a dashed curve), and again \( \rho = 0.4 \) with \( \eta = 0 \) (green triangles), and \( \rho = 0.3 \) with \( \eta = (\sqrt{5} - 1)/2 \) (red diamonds), for which we find \( 1/n \) decay (indicated by another dashed curve). The upper panel shows the data in semi-log, while the lower panel in log-log scale.
We note that \( \langle v_r \rangle = 0 \), so the decay of \( C(n) \) is an indicator of mixing. It is seen that, as expected, for \( \rho = 0.3 < \rho_c \), the correlation function \( C(n) \) decays algebraically, as \( \sim 1/n \). Quite obviously, the same algebraic decay takes place for the case with no magnetic field (\( \eta = 0 \), i.e., infinite-horizon Lorentz gas) independently on the disc radius. Instead, for \( \rho = 0.4 > \rho_c \) and for \( \eta < 1 \), the decay of \( C(n) \) is clearly exponential \( \langle C(n) \rangle \sim \exp(-\nu n) \).

The rate of exponential decay \( \nu \approx 0.45 \) appear not to sensibly depend on the magnetic field parameter \( \eta \). Analogously, the Lyapunov exponent, which is around \( \lambda = 1 \), only slightly depends on the value of \( \eta \).

We note that in our example of modified Sinai billiard, just like in the usual Sinai billiard, the central disk is relevant for introducing exponential instability into orbits’ dynamics. We have also studied a similar magnetic-mesh heterostructure without circular obstacles. Even though such a dynamics cannot produce Lyapunov instability it still makes dynamics nonseparable and nonintegrable for typical values of parameter \( \eta \). On the basis of numerical results we argue that such dynamics should be qualitatively similar to the dynamics of polygonal billiards. In particular, for our Cartesian magnetic mesh the rate of ergodicity seems very similar to the triangular billiards with two angles being irrational multiples of \( \pi \) while the third angle is a rational multiple of \( \pi \) [11]. Indeed, our magnetic periodic boundary conditions produce discontinuities in the Poincaré map—namely trajectories which are only slightly different can either penetrate or reflect from the boundary. Similar discontinuities are produced by scattering from nonright angles in polygonal billiards.

**Conclusion.**—We believe we provided the first example of a billiard type dynamical system with (piecewise) straight trajectories which breaks time reversal invariance and is ergodic and mixing with possibly exponential decay of correlations. The breaking of time reversal symmetry is known to have important consequences in several fields of physics. One important example is the generation of directed transport. Very recently it has been shown that it can play a crucial role in the design of high performance thermoelectric devices for efficient power generation and refrigeration [12]. In this relation, time irreversible billiards are the natural candidates to provide a better understanding of the dynamical mechanisms which determine the fundamental limits that thermodynamics imposes on thermal machines, a central issue in physics and of increasing relevance in the future society.

We acknowledge useful discussion with Carl Dettmann. The work has been supported by Grants No. P1-044 and No. J1-2208 of the Slovenian Research Agency (ARRS), MIUR-PRIN 2008, and by Regione Lombardia.

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